

Precision measurement of the positron fraction with the Alpha Magnetic Spectrometer

**A.Kounine / MIT
on behalf of the
AMS Collaboration**

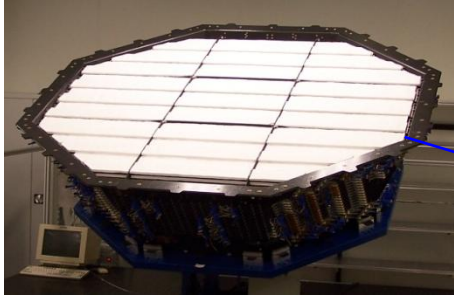


ICRC-2013, 8 July 2013, Rio de Janeiro

AMS: A TeV precision, multipurpose spectrometer in space.

TRD

Identify e^+ , e^-



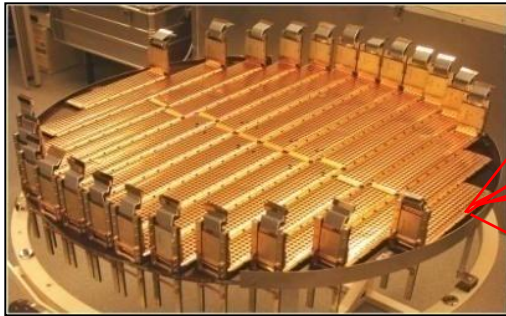
Positron fraction:

$$Ne^+ / (Ne^+ + Ne^-)$$

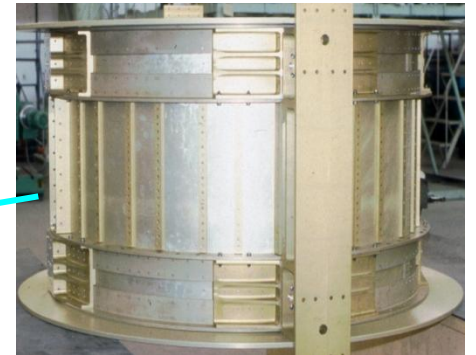
TOF
 Z, E



Silicon Tracker
 Z, P

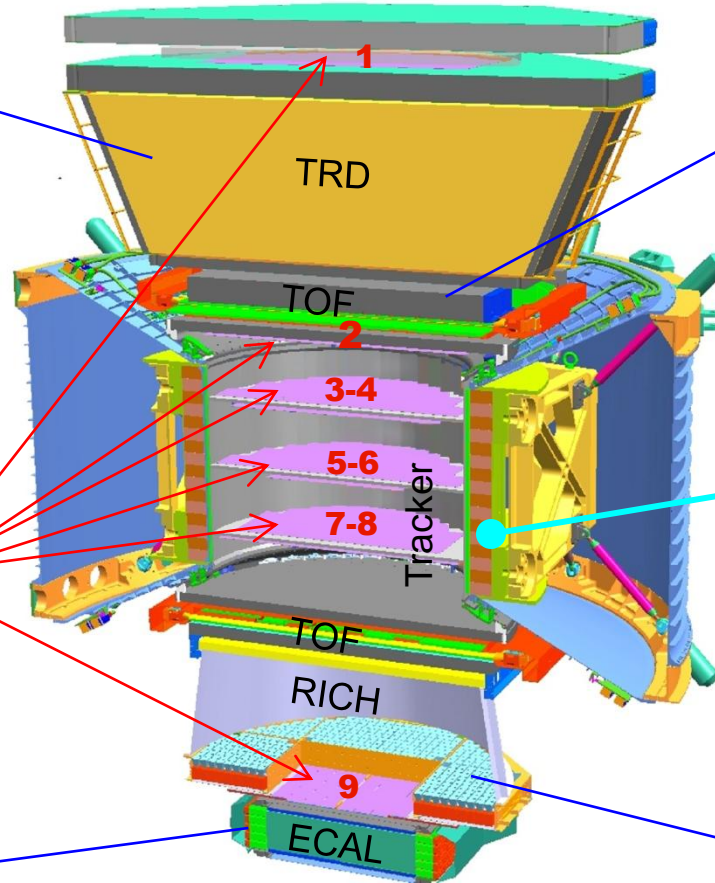


Magnet
 $\pm Z$

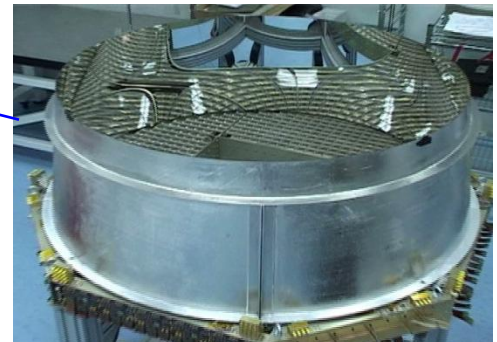


ECAL

E of e^+ , e^- , γ



RICH
 Z, E



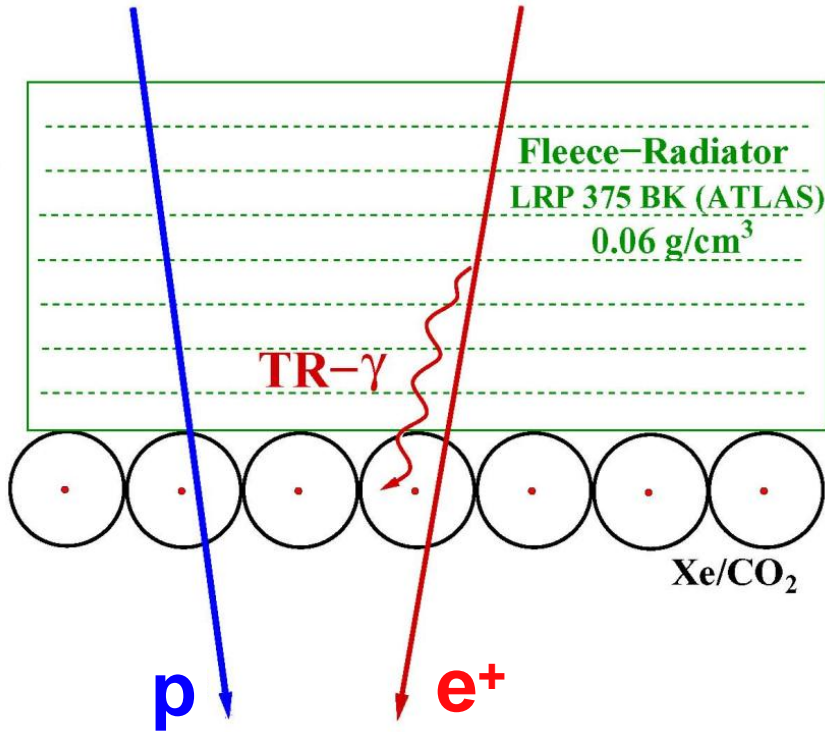
*Z, P are measured independently from
Tracker, RICH, TOF and ECAL*



Transition Radiation Detector.

Identify e^+ , reject P

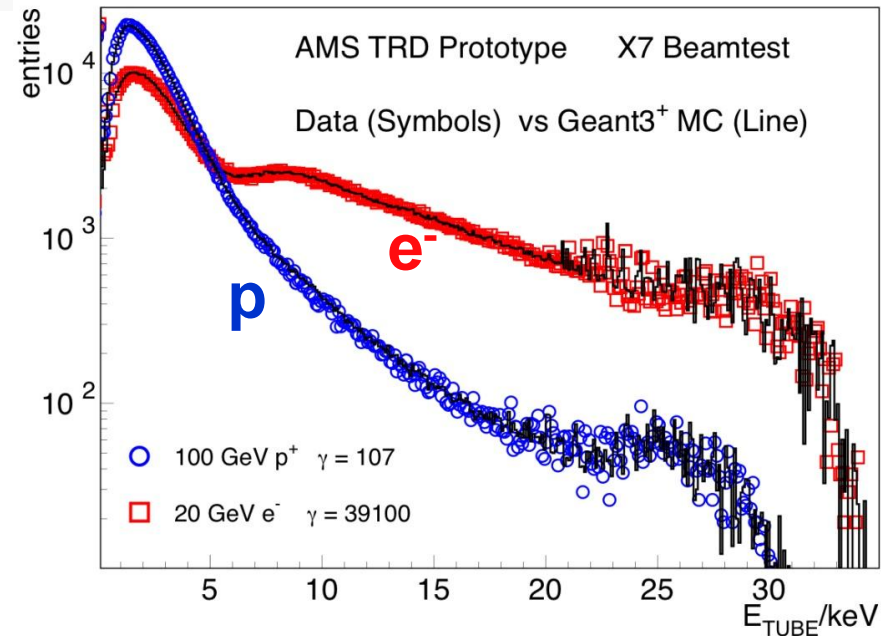
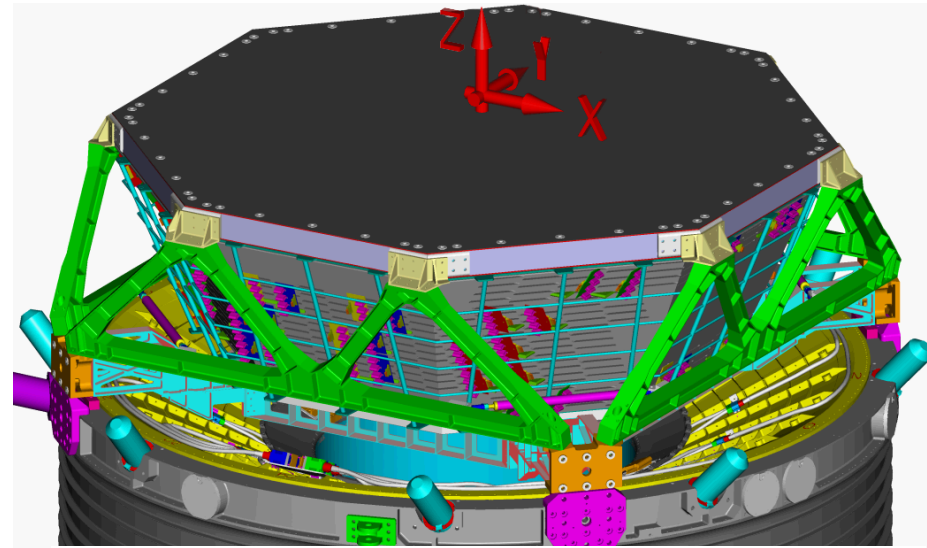
One of 20 Layers



Leak rate: CO₂ ≈ 5 μg/s

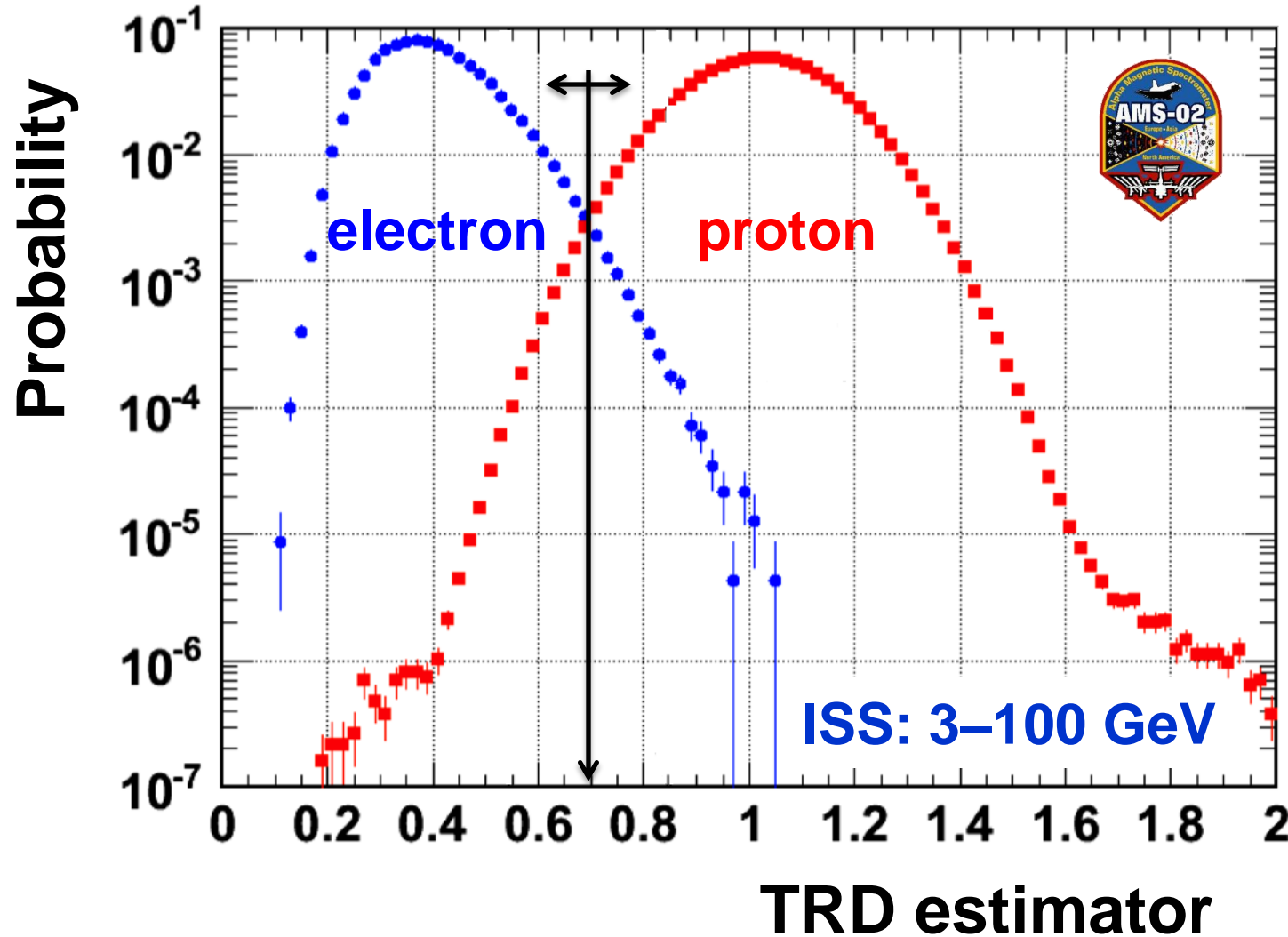
Storage: 5 kg, >20 years lifetime

Reported by H.Gast, ICRC



TRD performance on ISS

$$\text{TRD estimator} = -\ln(P_e / (P_e + P_p))$$



Normalized probabilities
 P_e and P_p

$$P_e = \sqrt[n]{\prod_i P_e^{(i)}(A)}$$

$$P_p = \sqrt[n]{\prod_i P_p^{(i)}(A)}$$

TRD performance on ISS

Proton rejection at 90% e^+ efficiency



• ISS data

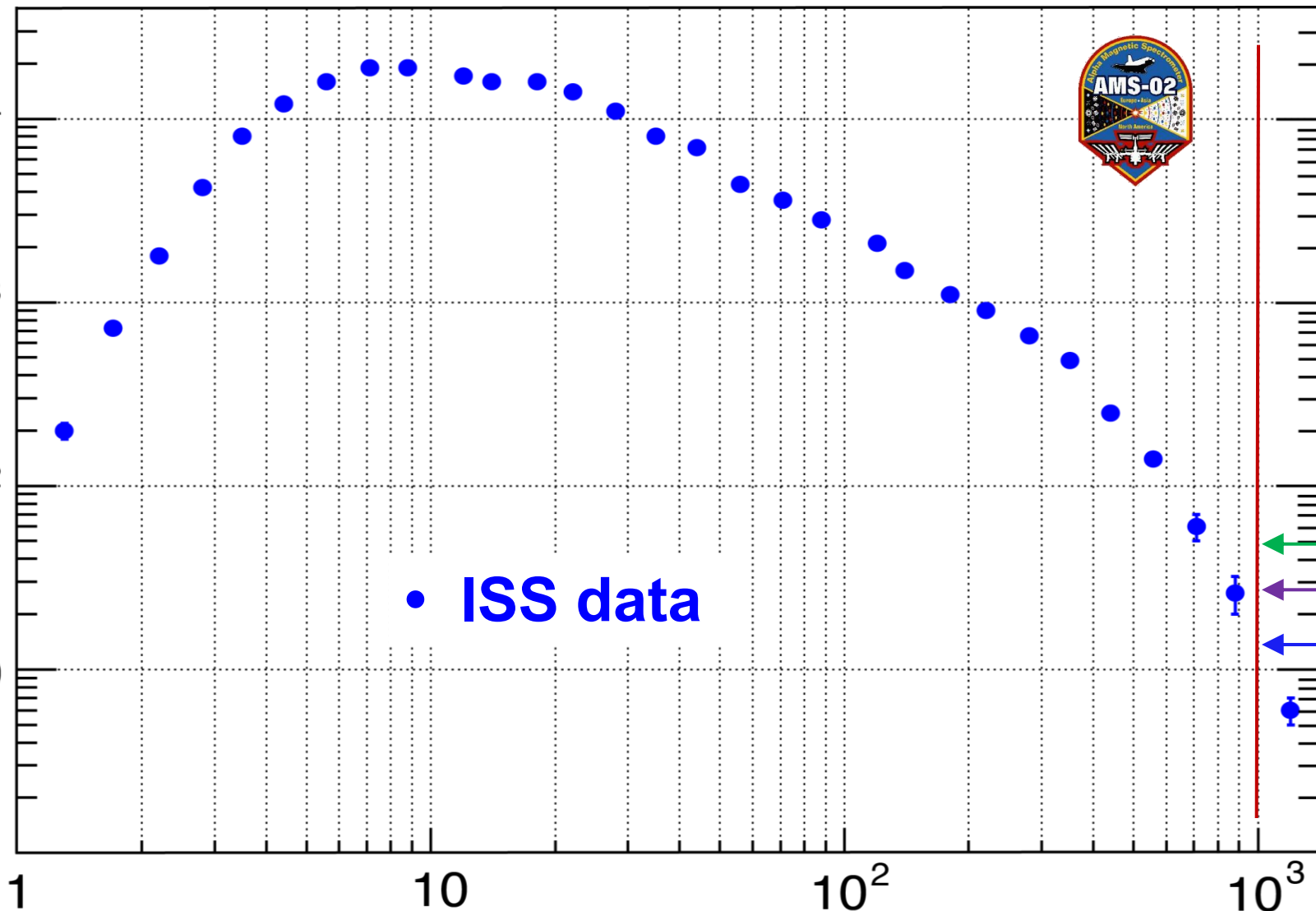
ϵ_e

70%

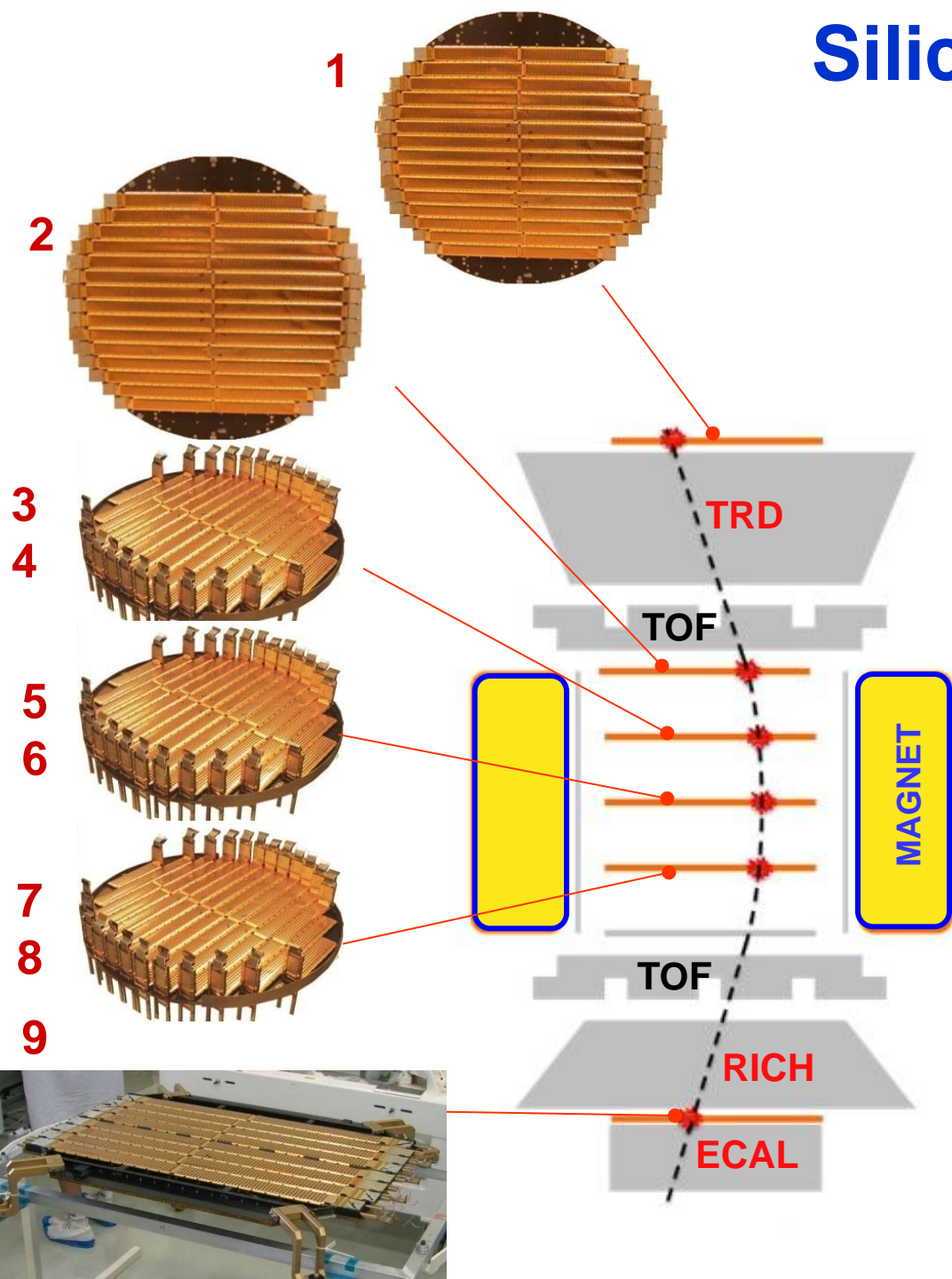
80%

90%

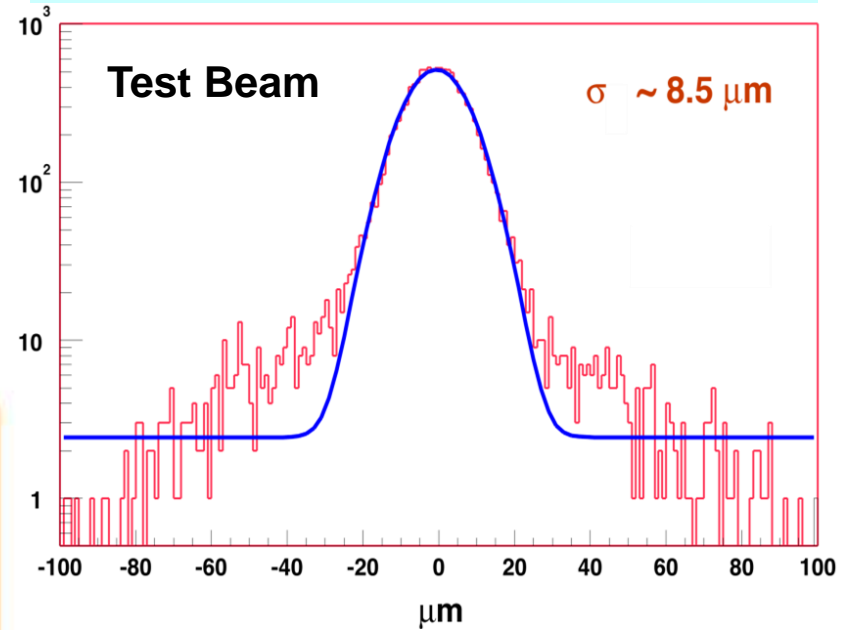
Rigidity (GV)



Silicon Tracker



Reported by P. Zuccon and
by C. Delgado, ICRC



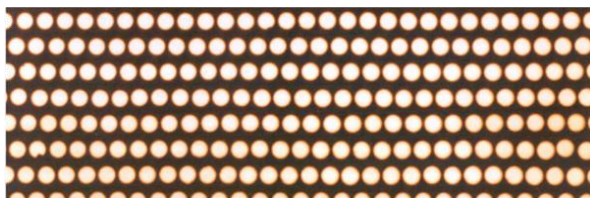
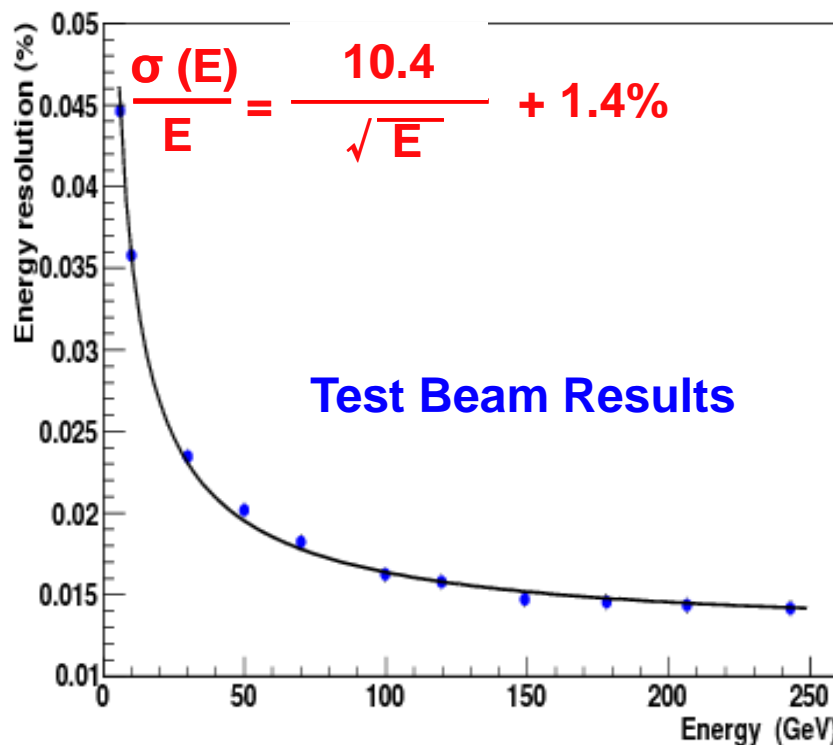
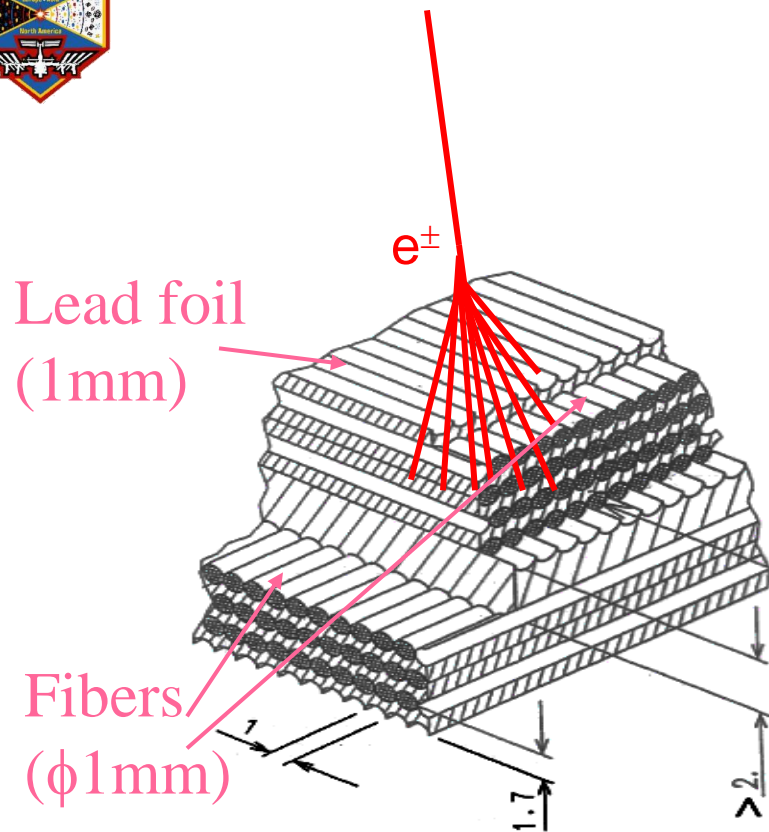
MDR $\sim 2.0 \text{ TV}$

$E / |p|$ matching



Electromagnetic Calorimeter

A precision, $17 X_0$, TeV, 3D measurement of the directions and energies of light rays and electrons



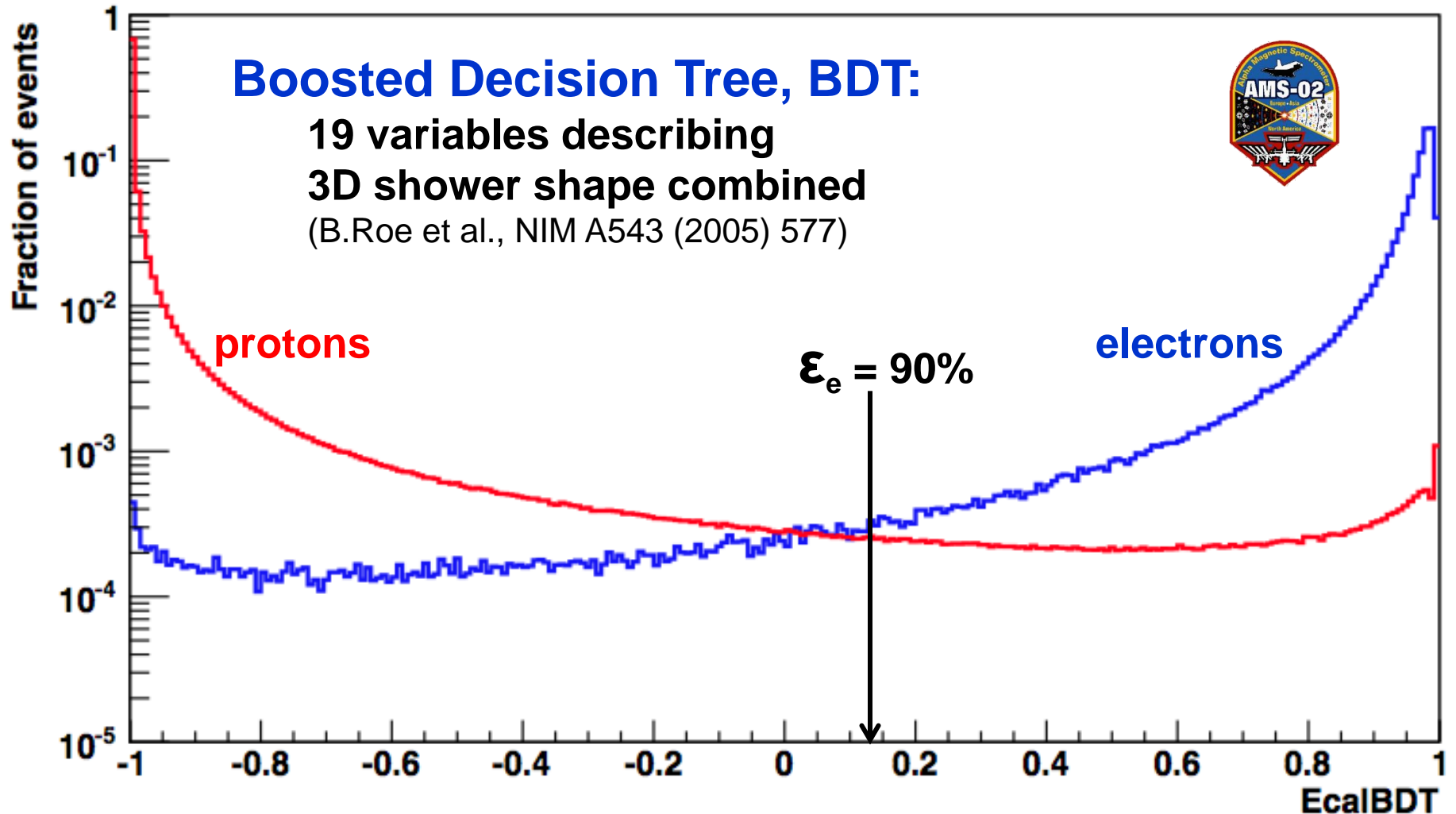
50 000 fibers, $\phi = 1\text{ mm}$
distributed uniformly
Inside 1,200 lb of lead



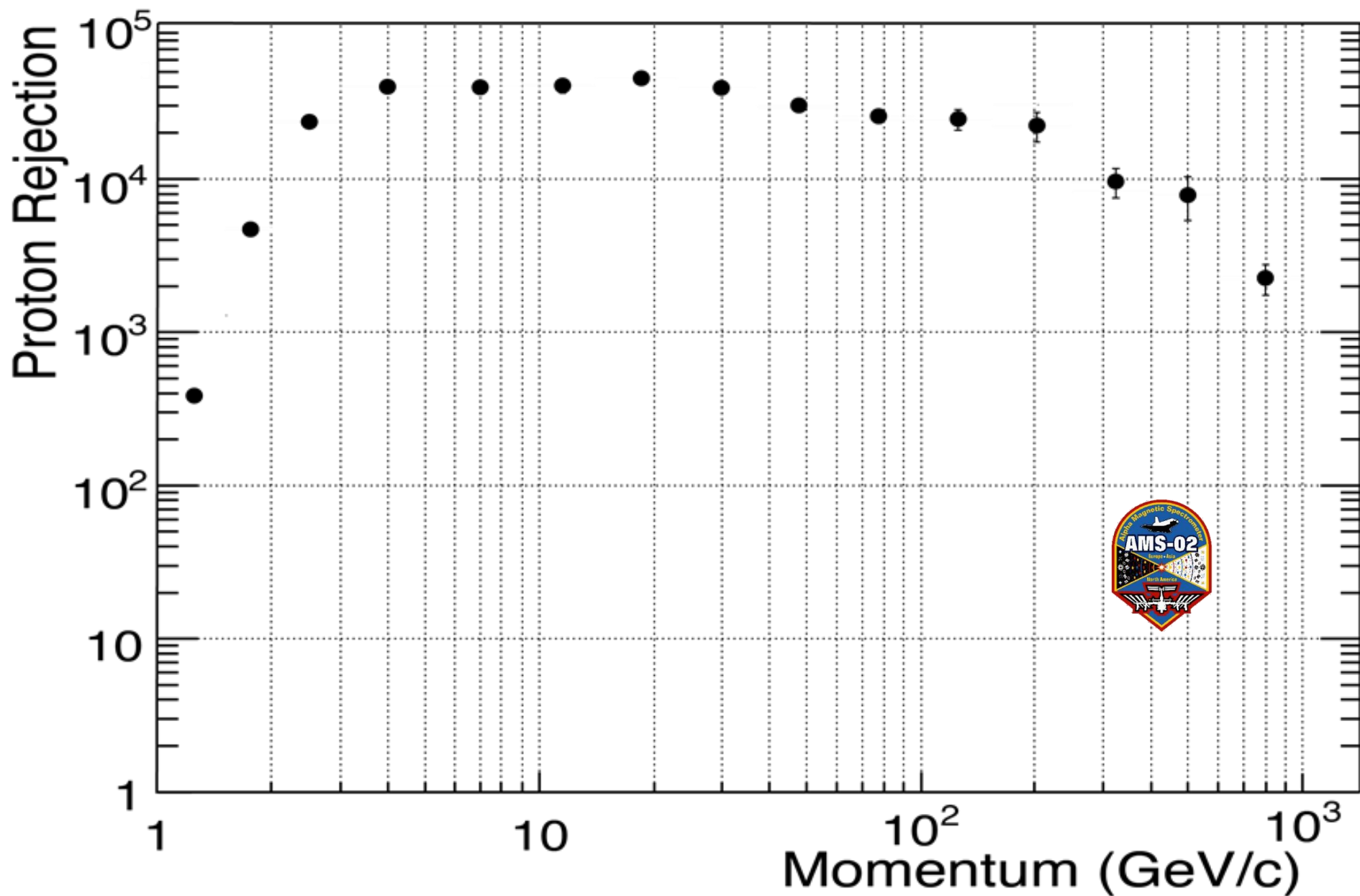
Reported by S. Di Falco, ICRC

Separation of protons and electrons with ECAL

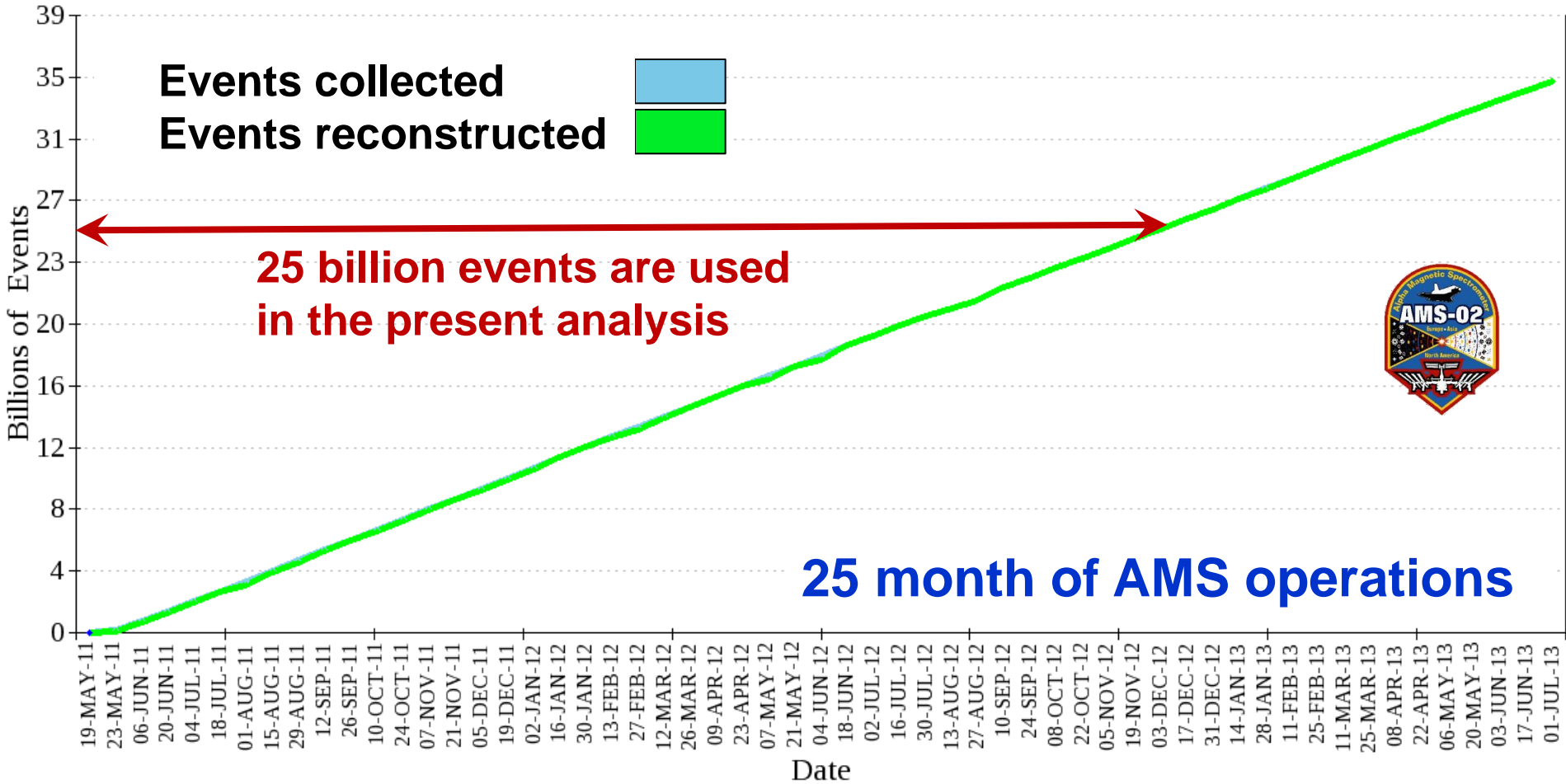
ISS data: 83–100 GeV



Data from ISS: Proton rejection using the ECAL

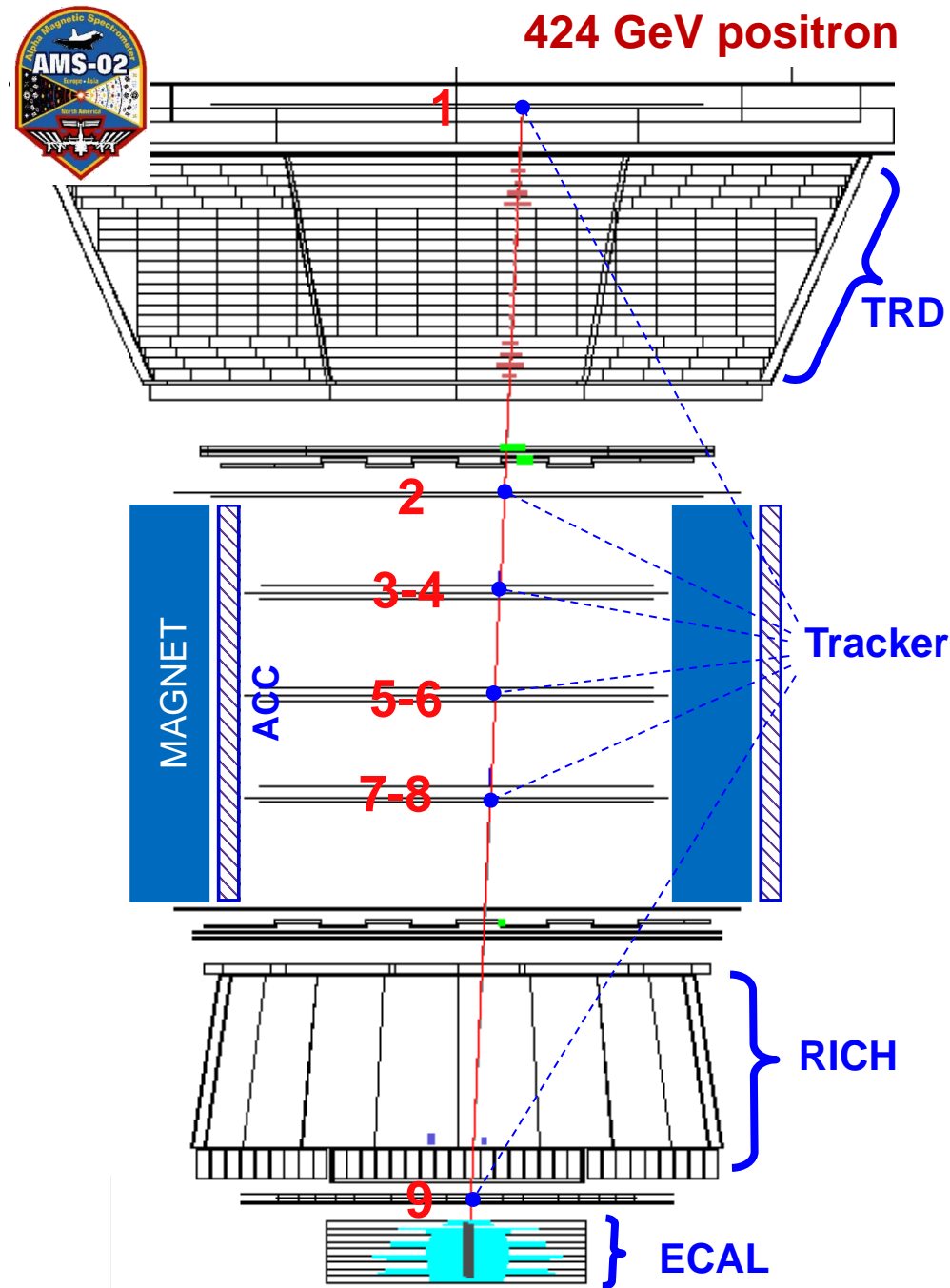


To date AMS collected over 35 billion events



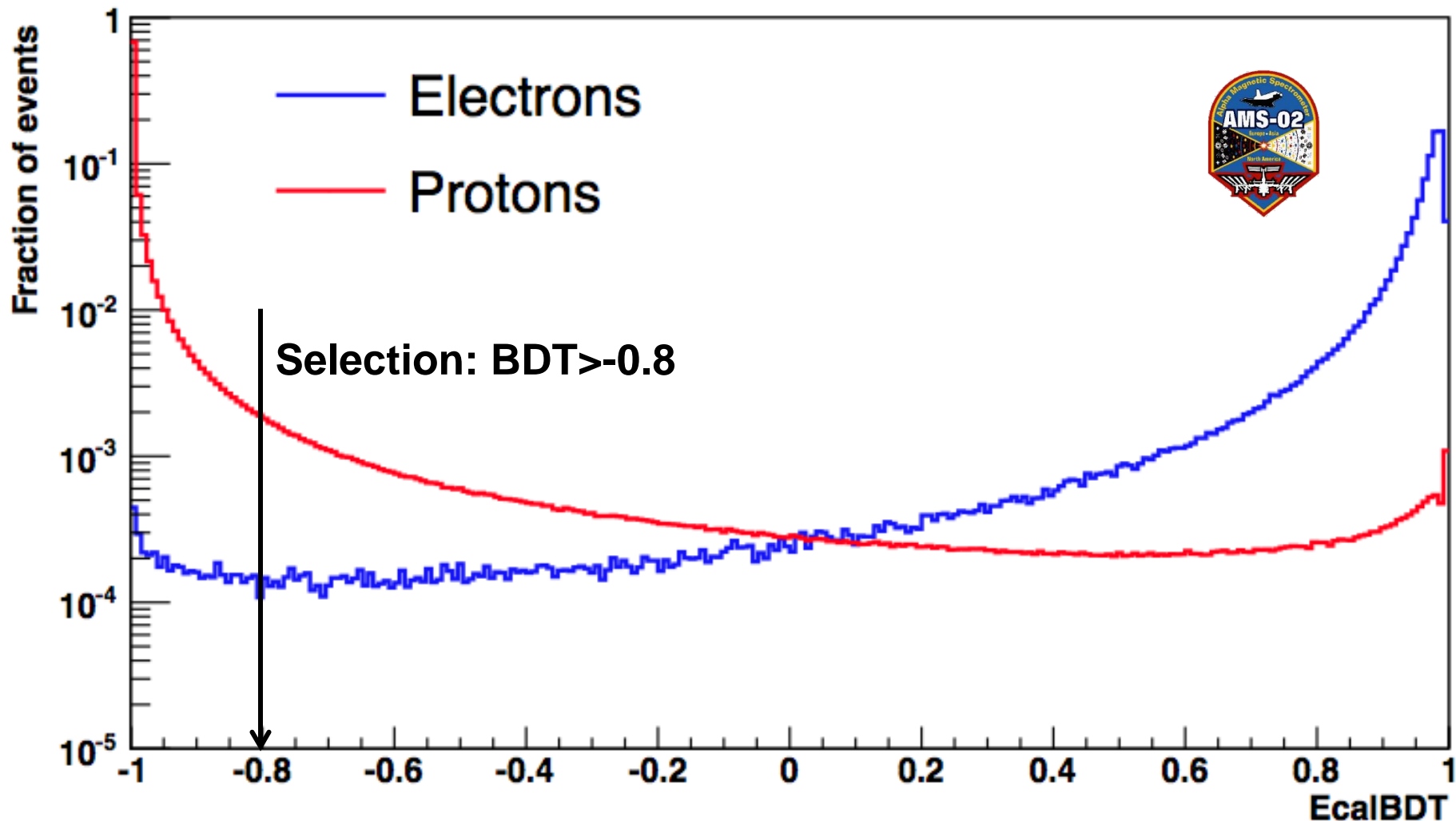
Event selection.

- **DAQ:** efficiency > 50% (no SAA)
- **Geomagnetic cutoff:**
 $E > 1.2 \cdot \text{max cutoff}$
- **TRACKER:**
 - Track quality
 - geometrical match with ECAL shower
- **TRD:** at least 15 hits
- **TOF:** downgoing particle,
 $\beta > 0.8$, $0.8 < Z < 1.4$
- **ECAL:**
 - shower axis within the fiducial ECAL volume
 - electromagnetic shape of the shower

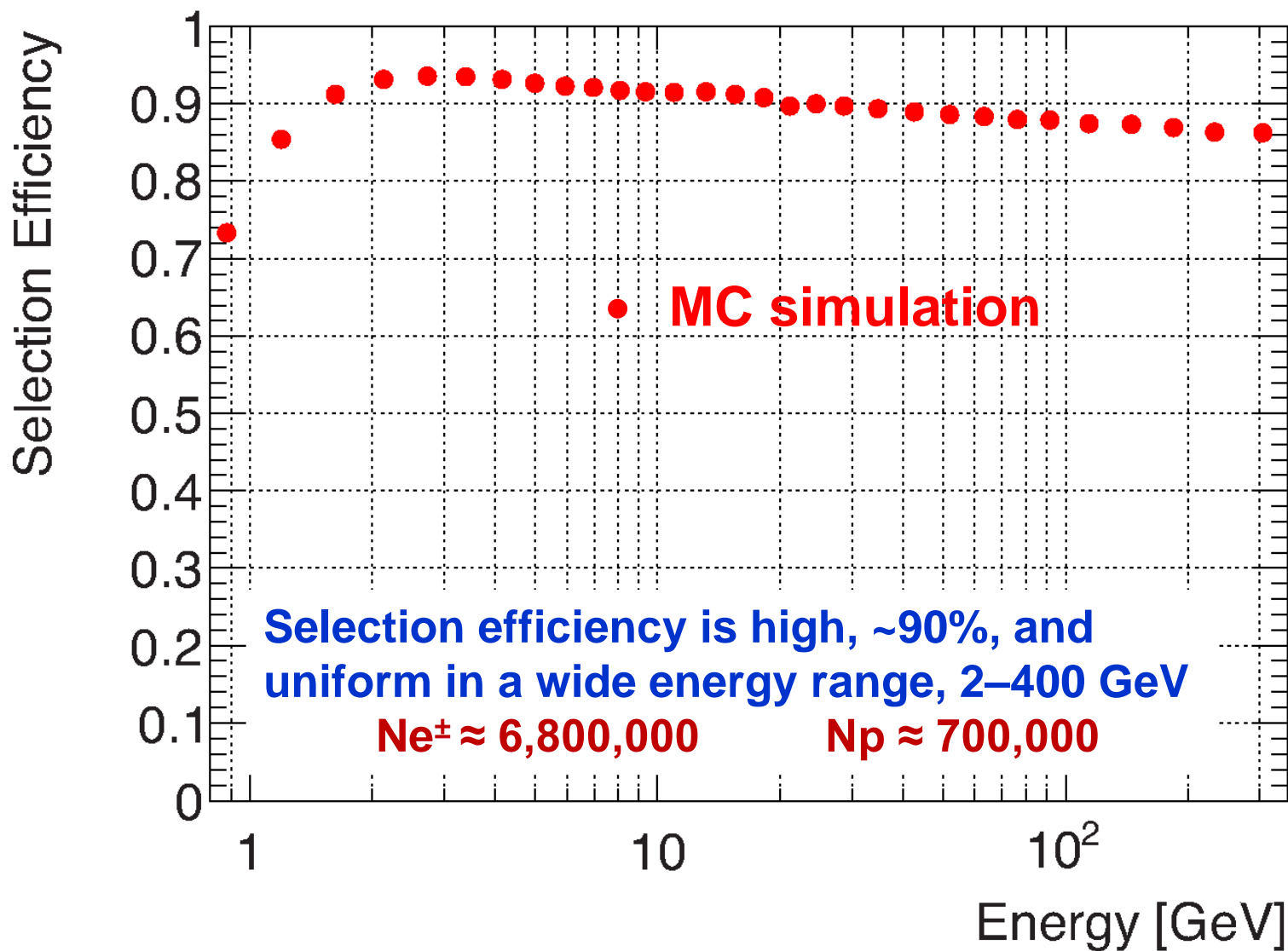


Event selection: ECAL BDT

ISS data: 83–100 GeV



Selection efficiency

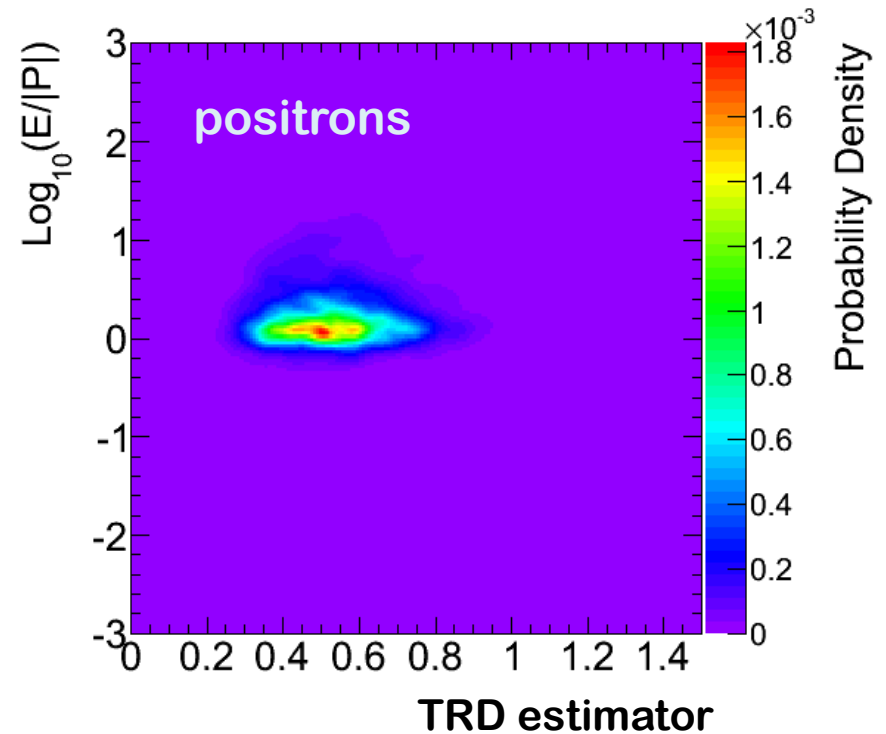
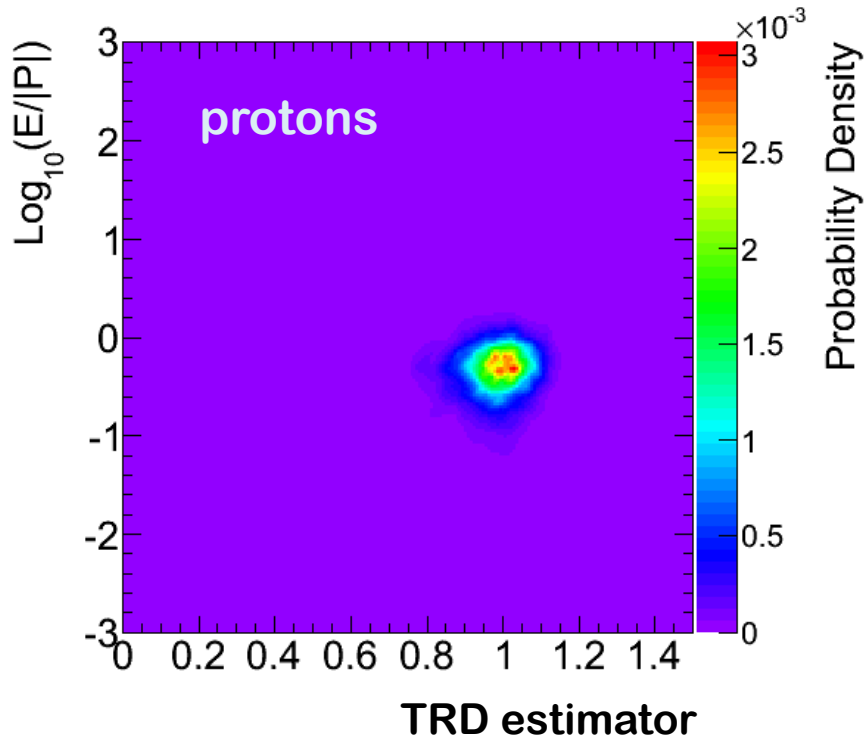




Analysis: 2D fit to measure $N_{e^{\pm}}$ and N_p

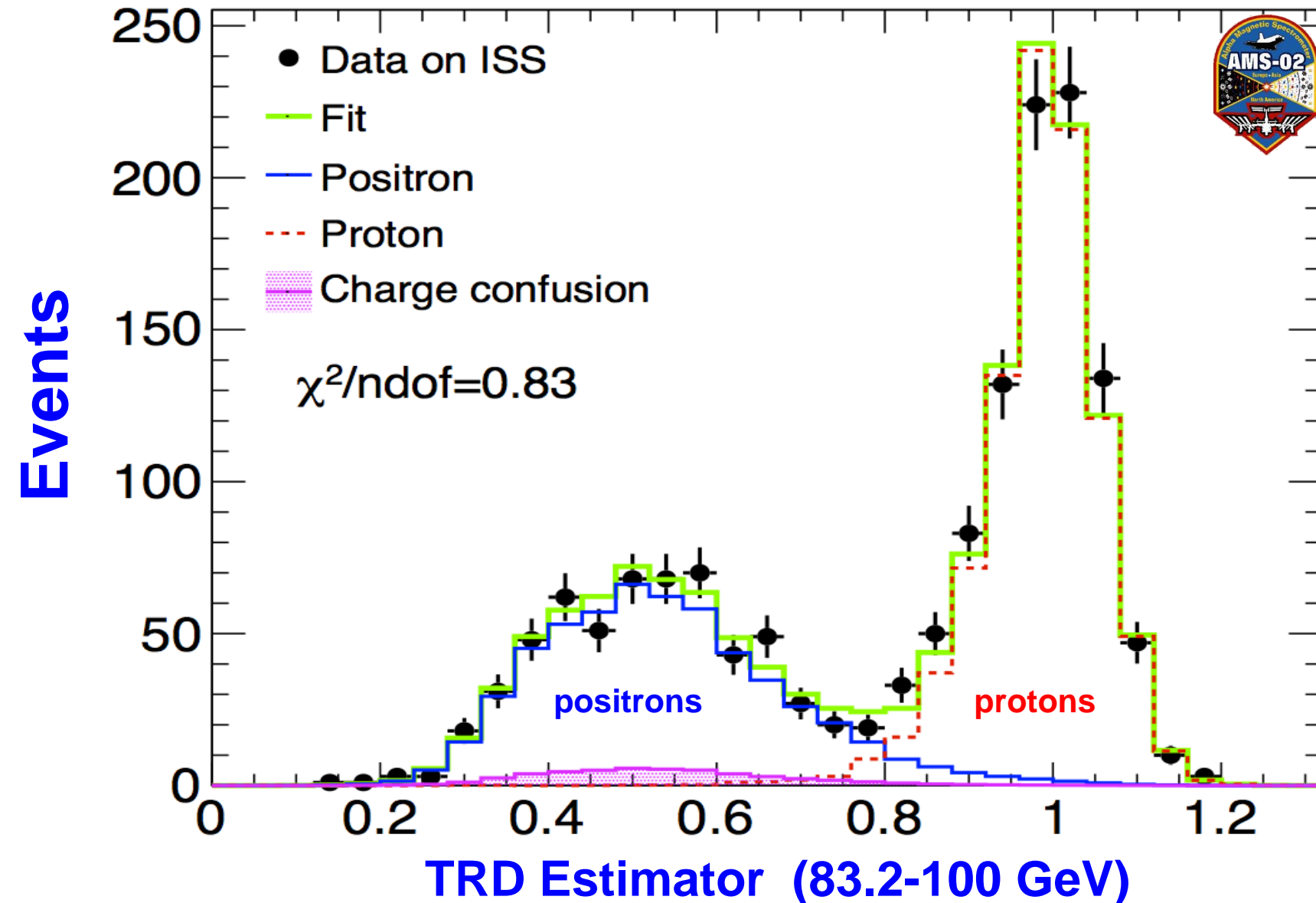
2D reference spectra for the signal and the background are fitted to data in the [TRD estimator- $\log(E/|P|)$] plane.

The method combines redundant information from TRD, ECAL, and Tracker; and provides much better statistical accuracy compared to cut-based analysis.

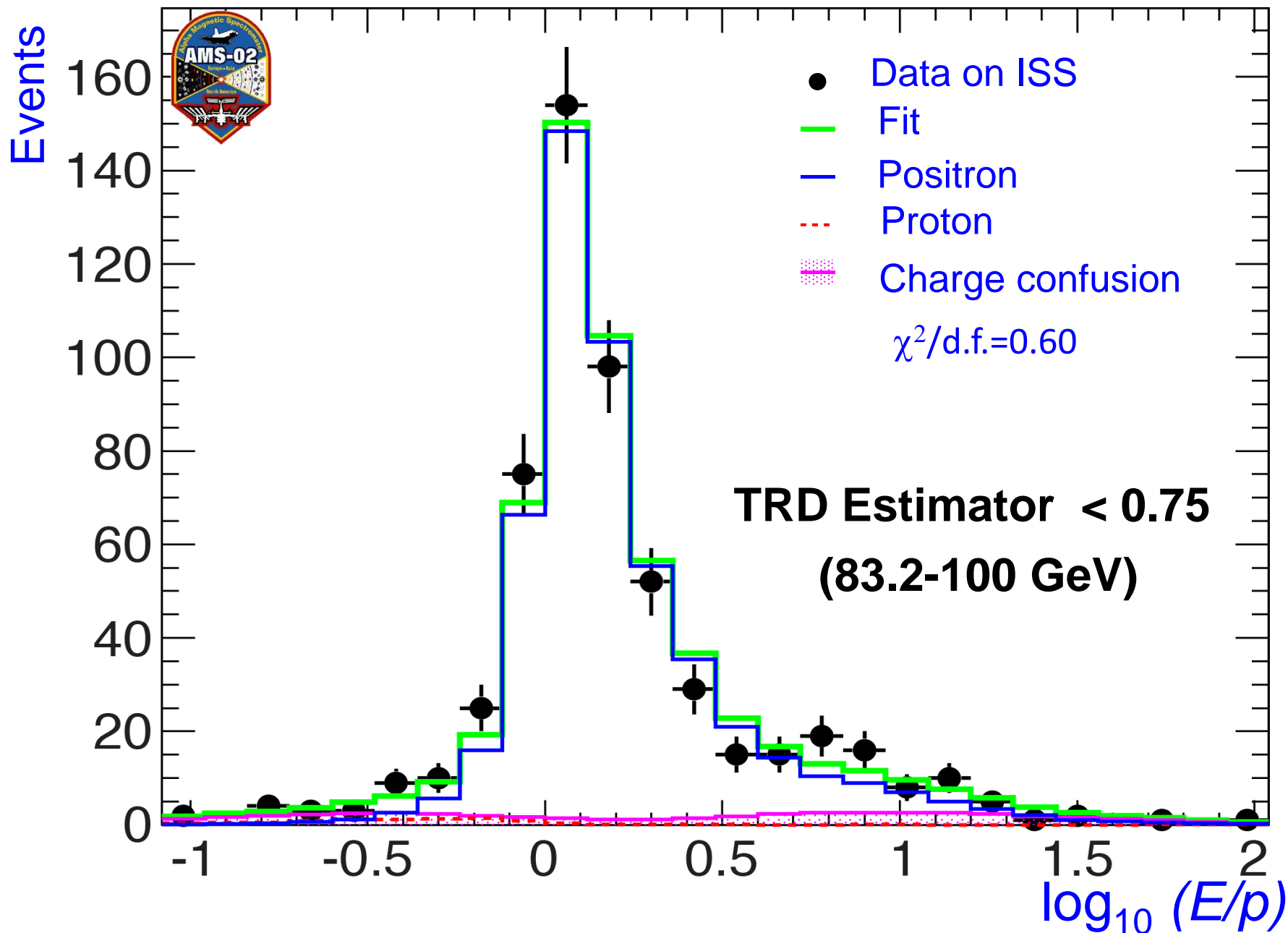


Results of the fit:

The TRD Estimator shows clear separation between **protons** and positrons with a small **charge confusion** background

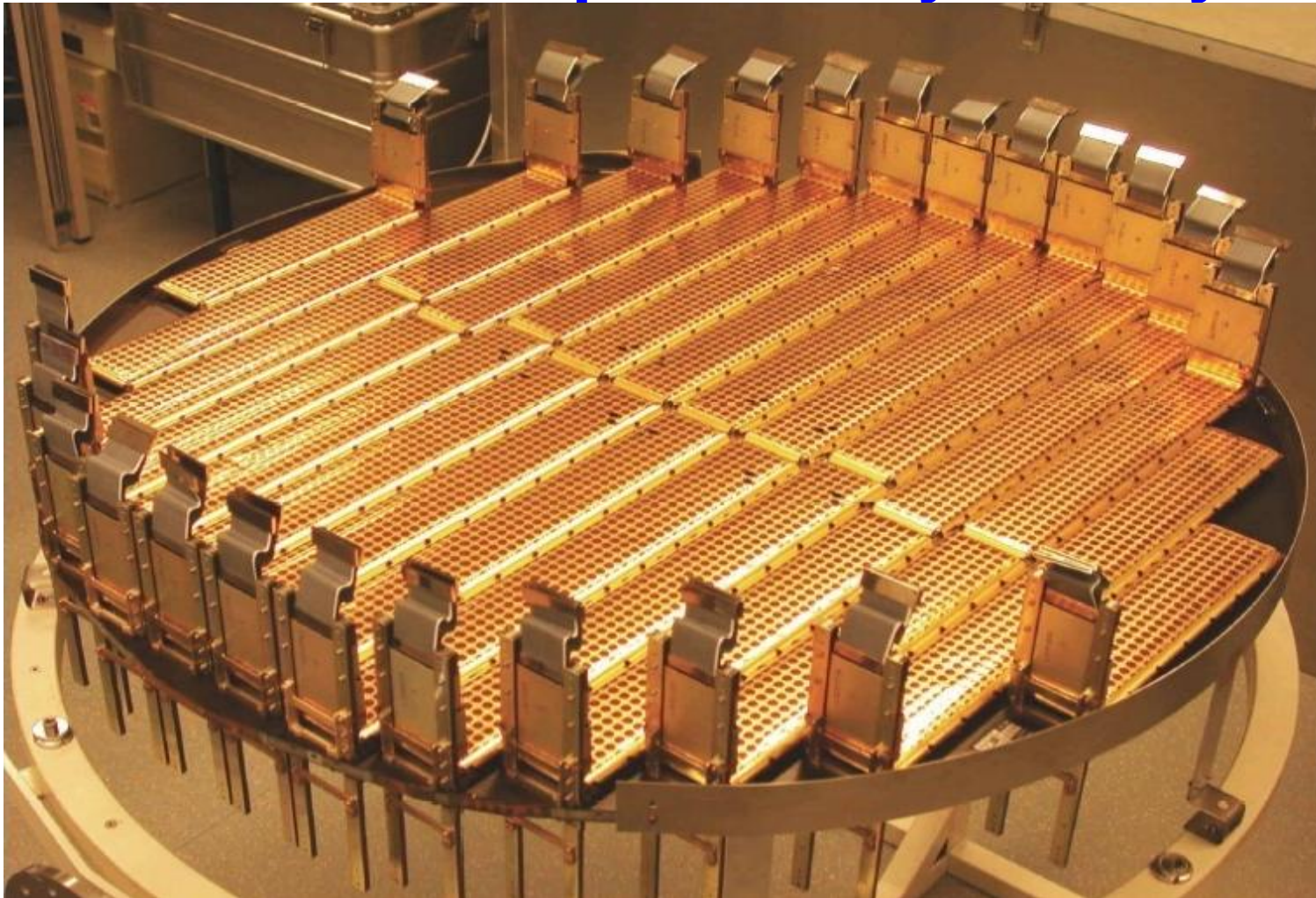


Results of the fit: in the signal region only 1 % of protons





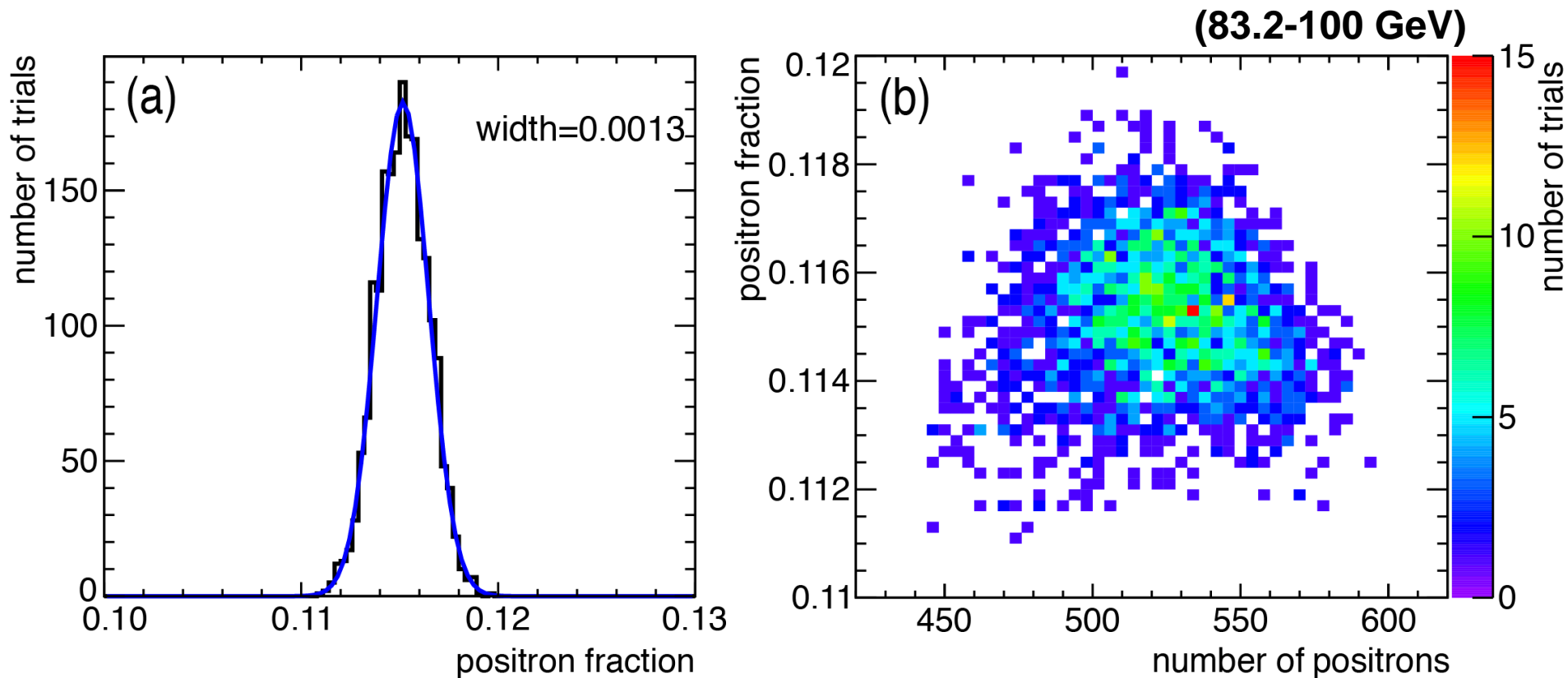
Systematic error on the positron fraction: 1. acceptance asymmetry



Difference between positron and electron acceptance
due to known minute tracker asymmetry



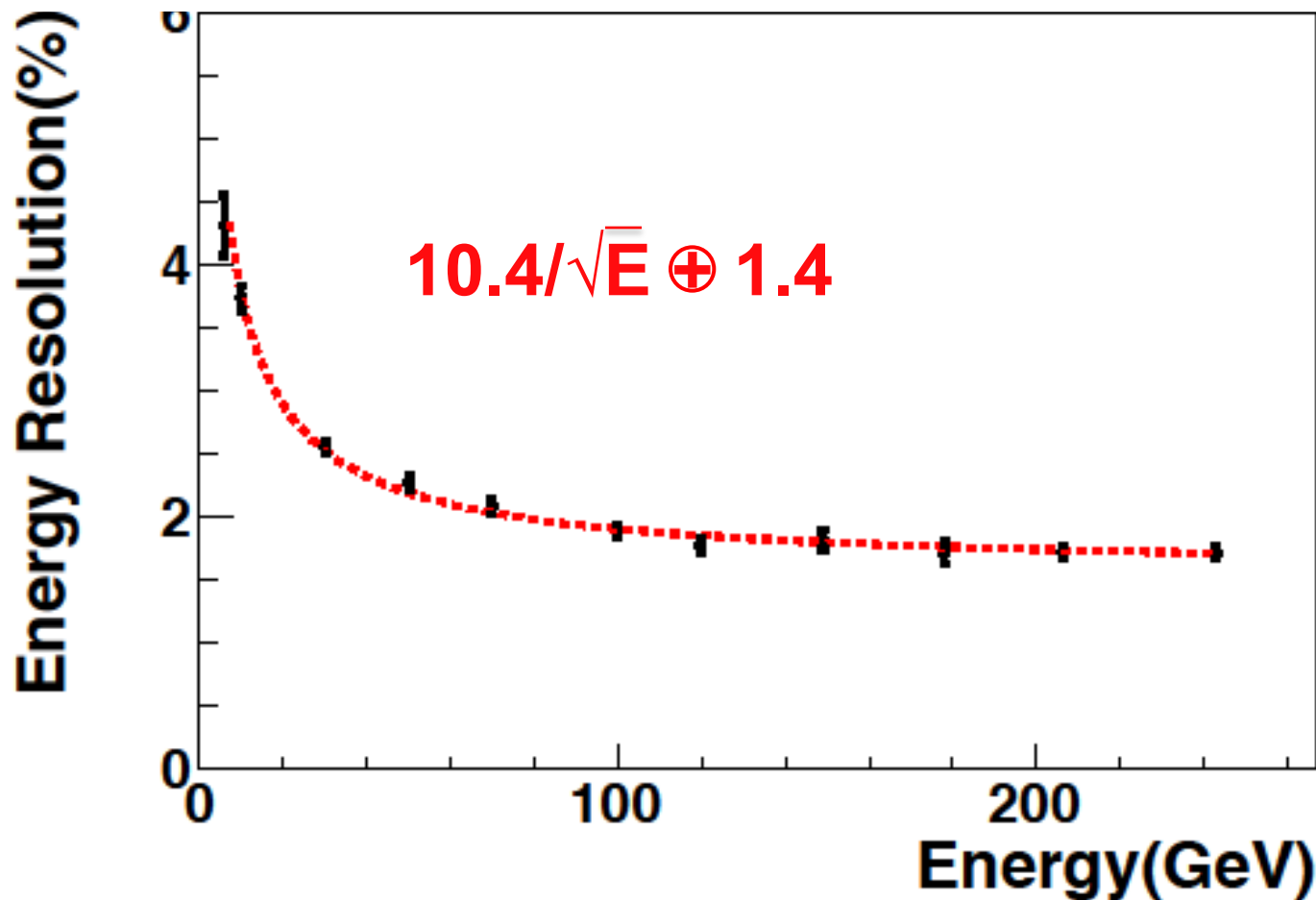
Systematic error on the positron fraction: 2. Selection dependence



The measurement is stable over wide variations of the cuts in the TRD identification, ECAL Shower Shape, E (from ECAL) matched to $|P|$ (from the Tracker), ... For each energy bin, over 1,000 sets of cuts were analyzed.



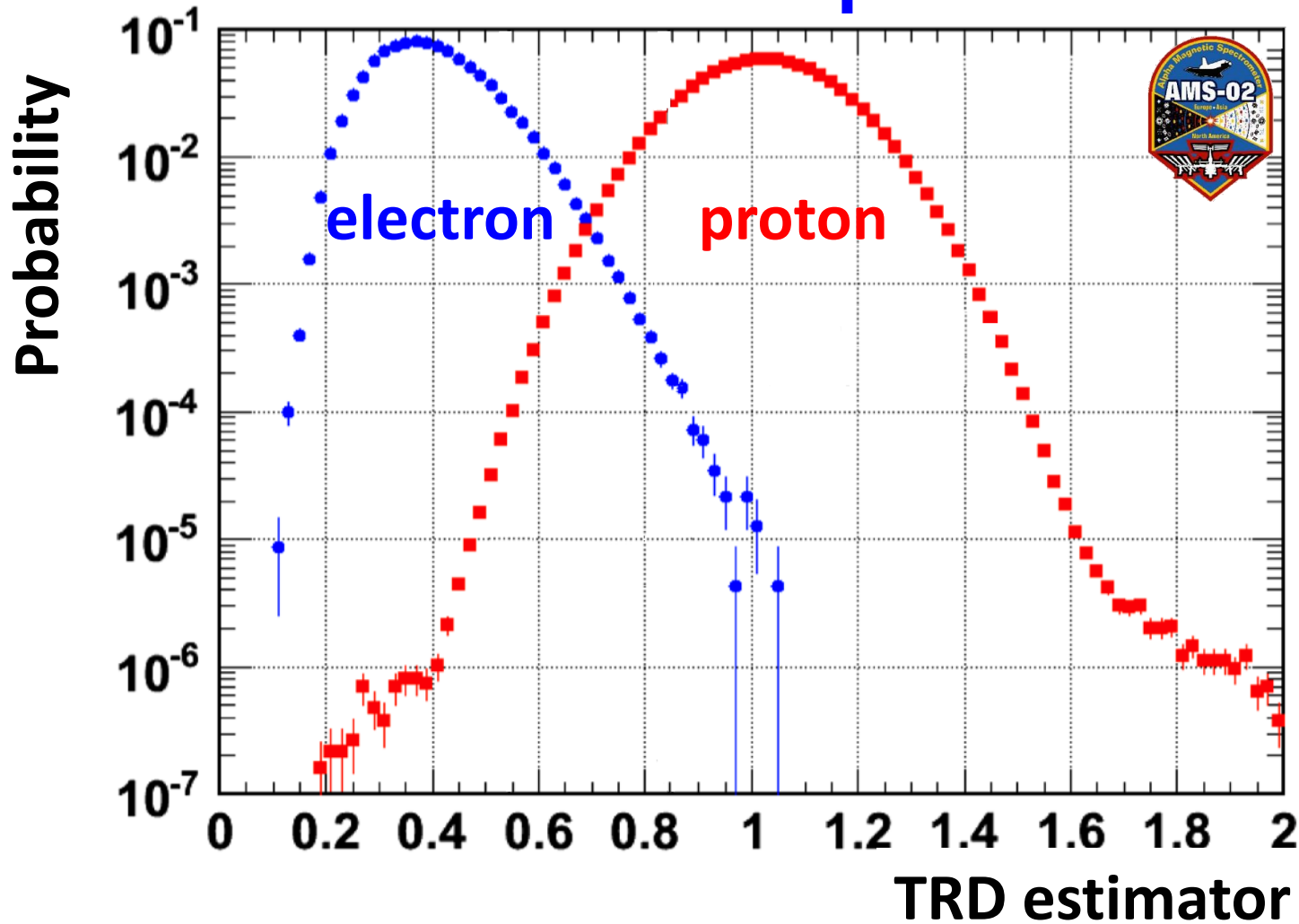
Systematic error on the positron fraction: 3. Bin-to-bin migration



Event migration effects are obtained by folding the measured spectra of positrons and electrons with the ECAL energy resolution.

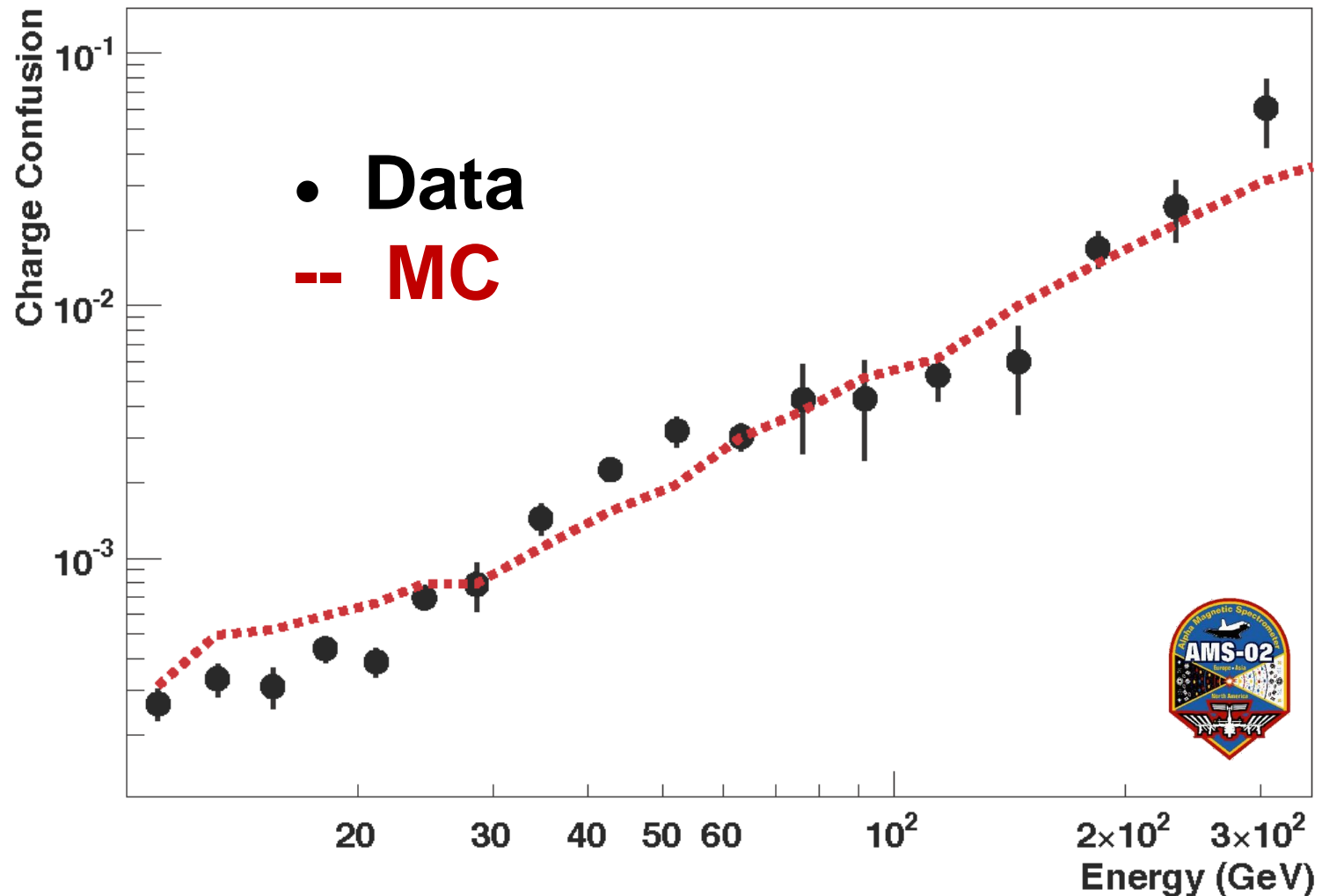
Bin width: 2σ at 5 GeV; 4σ at 50 GeV; 8σ at 100 GeV; 19σ at 300 GeV.

Systematic error on the positron fraction: 4. Reference spectra



Definition of the reference spectra is based on pure samples of electrons and protons of finite statistics.

Systematic error on the positron fraction: 5. e+/- Charge confusion



Two sources: large angle scattering and production of secondary tracks along the path of the primary track. Both are well reproduced by MC. Systematic errors correspond to variations of these effects within their statistical limits.

AMS Result: Measurement of the positron fraction



Positron events, positron fraction in each energy bin

Systematic Errors

Energy [GeV]	N_{e^+}	Fraction	statistical error	acceptance asymmetry	event selection	bin-to-bin migration	reference spectra	charge confusion	total systematic uncertainty
Energy[GeV]	N_{e^+}	Fraction	$\sigma_{stat.}$	$\sigma_{acc.}$	$\sigma_{sel.}$	$\sigma_{mig.}$	$\sigma_{ref.}$	$\sigma_{c.c.}$	$\sigma_{syst.}$
1.00-1.21	9335	0.0842	0.0008	0.0005	0.0009	0.0008	0.0001	0.0005	0.0014
1.97-2.28	23893	0.0642	0.0004	0.0002	0.0005	0.0002	0.0001	0.0002	0.0006
3.30-3.70	20707	0.0550	0.0004	0.0001	0.0003	0.0000	0.0001	0.0002	0.0004
6.56-7.16	13153	0.0510	0.0004	0.0001	0.0000	0.0000	0.0001	0.0002	0.0002
09.95-10.73	7161	0.0519	0.0006	0.0001	0.0000	0.0000	0.0001	0.0002	0.0002
19.37-20.54	2322	0.0634	0.0013	0.0001	0.0001	0.0000	0.0001	0.0002	0.0003
30.45-32.10	1094	0.0701	0.0022	0.0001	0.0002	0.0000	0.0001	0.0003	0.0004
40.00-43.39	976	0.0802	0.0026	0.0002	0.0005	0.0000	0.0001	0.0004	0.0007
50.87-54.98	605	0.0891	0.0038	0.0002	0.0006	0.0000	0.0001	0.0004	0.0008
64.03-69.00	392	0.0978	0.0050	0.0002	0.0010	0.0000	0.0002	0.0007	0.0013
74.30-80.00	276	0.0985	0.0062	0.0002	0.0010	0.0000	0.0002	0.0010	0.0014
86.00-92.50	240	0.1120	0.0075	0.0002	0.0010	0.0000	0.0003	0.0011	0.0015
100.0-115.1	304	0.1118	0.0066	0.0002	0.0015	0.0000	0.0003	0.0015	0.0022
115.1-132.1	223	0.1142	0.0080	0.0002	0.0019	0.0000	0.0004	0.0019	0.0027
132.1-151.5	156	0.1215	0.0100	0.0002	0.0021	0.0000	0.0005	0.0024	0.0032
151.5-173.5	144	0.1364	0.0121	0.0002	0.0026	0.0000	0.0006	0.0045	0.0052
173.5-206.0	134	0.1485	0.0133	0.0002	0.0031	0.0000	0.0009	0.0050	0.0060
206.0-260.0	101	0.1530	0.0160	0.0003	0.0031	0.0000	0.0013	0.0095	0.0101
260.0-350.0	72	0.1550	0.0200	0.0003	0.0056	0.0000	0.0018	0.0140	0.0152

Positron fraction

The data show that the positron fraction is steadily increasing from 10 to ~250 GeV, but, from 20 to 250 GeV, the slope decreases by an order of magnitude.

- AMS-02 (6.8 million e^+ , e^- events)
8% of total Data to 2028

10^{-1}



No fine structure in the spectrum

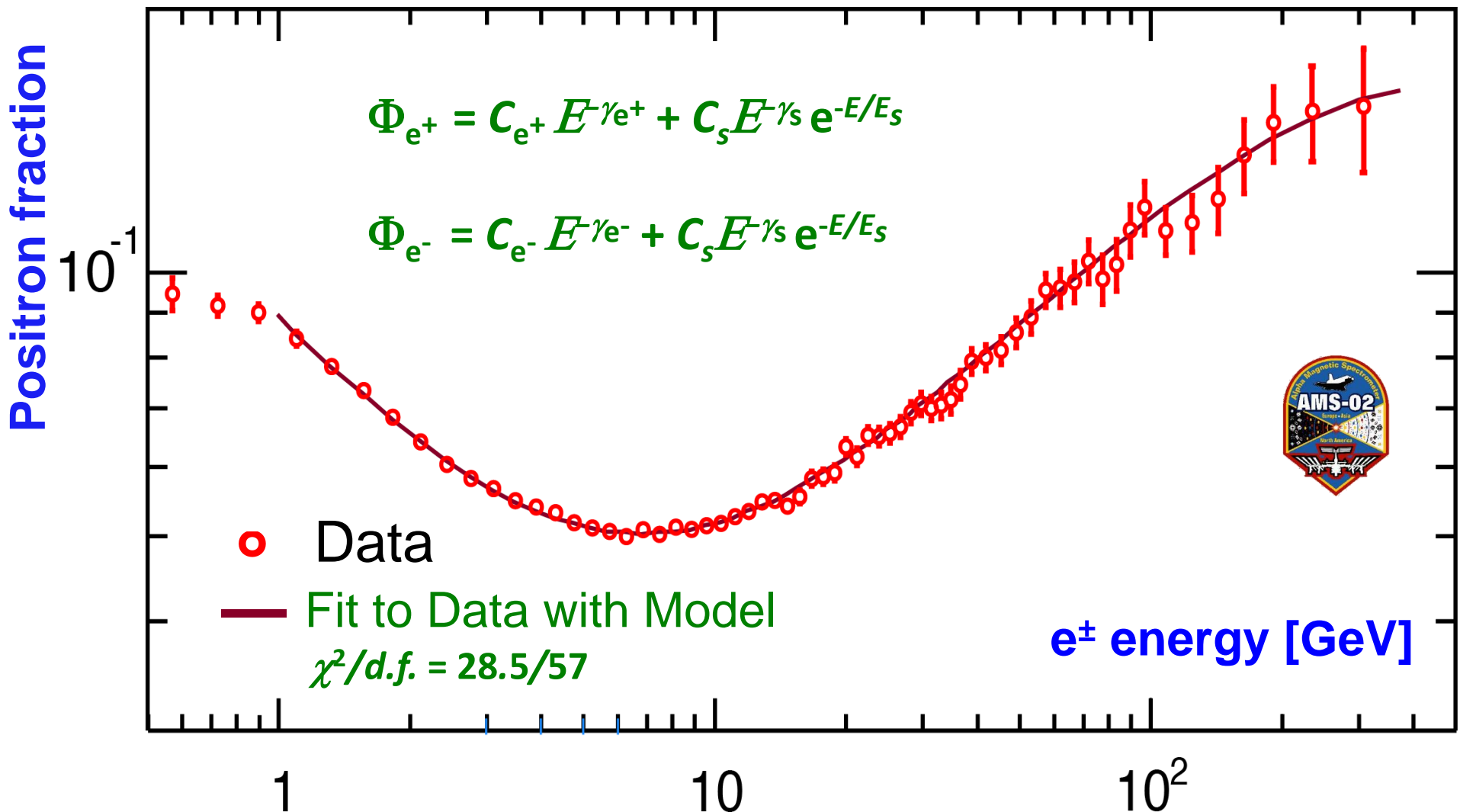
e^\pm energy [GeV]

1

10

10^2

Physics Example: Comparing data with a minimal model.



The agreement between the data and the model shows that the positron fraction spectrum is consistent with e^\pm fluxes each of which is the sum of its diffuse spectrum and a single common power law source.



A fit to the data in the energy range 1 to 350 GeV yields:

$\gamma_{e^-} - \gamma_{e^+} = -0.63 \pm 0.03$, *i.e.*, the diffuse positron spectrum is less energetic than the diffuse electron spectrum;

$\gamma_{e^-} - \gamma_s = 0.66 \pm 0.05$, *i.e.*, the source spectrum is more energetic than the diffuse electron spectrum;

$C_{e^+}/C_{e^-} = 0.091 \pm 0.001$, *i.e.*, the weight of the diffuse positron flux amounts to $\sim 10\%$ of that of the diffuse electron flux;

$C_s/C_{e^-} = 0.0078 \pm 0.0012$, *i.e.*, the weight of the common source constitutes only $\sim 1\%$ of that of the diffuse electron flux;

$1/E_s = 0.0013 \pm 0.0007 \text{ GeV}^{-1}$,
corresponding to a cutoff energy of $760^{+1000}_{-280} \text{ GeV}$.



Positron fraction

10^{-1}

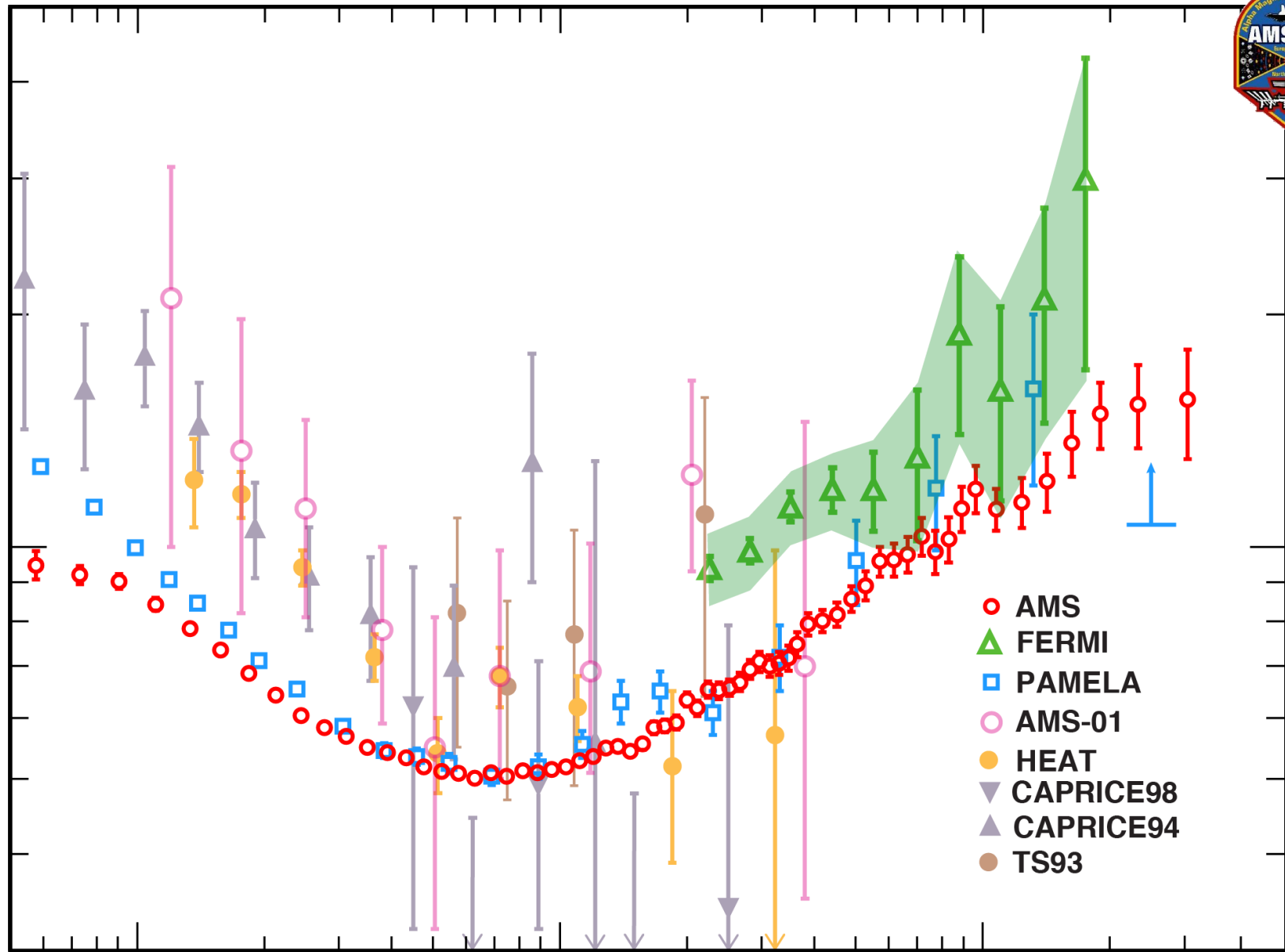
1

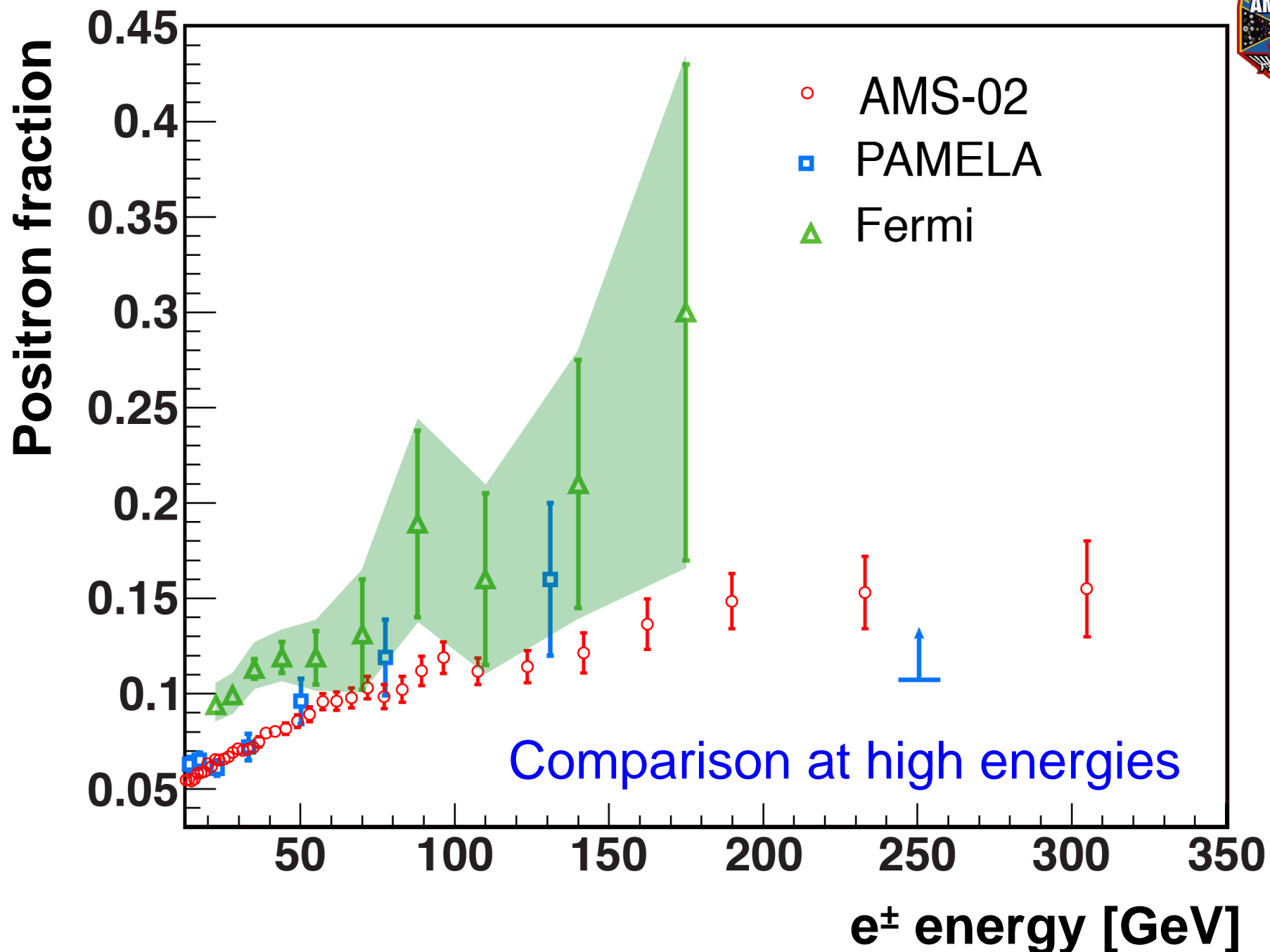
10

10^2

positron, electron energy [GeV]

- AMS
- △ FERMI
- PAMELA
- AMS-01
- HEAT
- ▼ CAPRICE98
- ▲ CAPRICE94
- TS93







**In conclusion,
the first 6.8 million primary positron and electron events collected
with AMS on the ISS show:**

- I. At energies < 10 GeV, a decrease in the positron fraction with increasing energy.
- II. A steady increase in the positron fraction from 10 to ~ 250 GeV.
- III. The determination of the behavior of the positron fraction from 250 to 350 GeV and beyond requires more statistics.
- IV. The slope of the positron fraction versus energy decreases by an order of magnitude from 20 to 250 GeV and no fine structure is observed. The agreement between the data and the model shows that the positron fraction spectrum is consistent with e^\pm fluxes each of which is the sum of its diffuse spectrum and a single common power law source.

**These observations show the existence of new physical phenomena,
whether from a particle physics or an astrophysical origin.**

Electron E=982 GeV

Run/Event 1329775818/ 60709

Positron E=636 GeV

Run/Event 133119-743/ 56950

